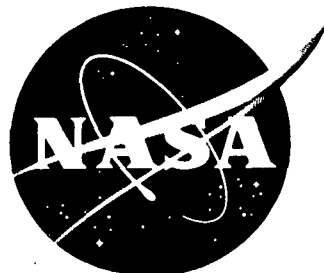


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NASA TM X-58101
November 1972



STRATIGRAPHY OF THE APOLLO 15 DRILL CORE

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16. Abstract <p>The crew of Apollo 15 collected a 242-centimeter-long core of the regolith of the moon developed on the surface of Palus Putredinis (3° 39'20"E, 26° 26'00"N). The 2.04-centimeter-diameter core, which has a mass of 1333.2 grams, consists of 42 major textural units, with thicknesses ranging from a few milliliters to 13 centimeters thick. The regolith is not homogeneous and is composed of many layers that are mostly ejecta from impact events.</p>					
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STRATIGRAPHY OF THE APOLLO 15 DRILL CORE

By Grant Heiken, * Michael Duke, * Roald Fryxell,[†]
J. Stewart Nagle,[‡] Ron Scott,[§] and G. A. Sellers^{||}

INTRODUCTION

The Hadley Rille-Apennine Mountains area of the moon ($3^{\circ} 39' 20'' \text{E}$, $26^{\circ} 26' 00'' \text{N}$) was explored by the crew of Apollo 15. Among the 78 kilograms of samples that were returned were one 36-centimeter- and two 64-centimeter-long hand-driven cores and one 242-centimeter-long core collected with a rotary-percussion electric drill. The 242-centimeter-long core, which was collected from the regolith developed on Palus Putredinis, is hopefully a representative section of the regolith developed on the mare surface, although its location at station 8, 50 meters from the Apollo lunar surface experiments package (ALSEP) (fig. 1, from ref. 1) central station, may have been on the edge of a ray (ref. 2).

The purpose of this report is to provide a stratigraphic description based on visual observations made during the dissection of the drill core. This description should serve as a basis on which detailed studies may be carried out.

Those responsible for description and dissection of the core sections are as follows: 15001, Fryxell and Heiken; 15002, Duke and Heiken; 15003, Nagle; 15004, Heiken and Scott; 15005, Sellers and Heiken; and 15006, Heiken and Fryxell. All were ably assisted by Maureen Mitchell, Ed Cornetius, or R. White. Henry Cantu operated the X-ray unit, and Paul Gilmore or Al Locke were responsible for the photographs.

EQUIPMENT AND SAMPLING METHOD

A lightweight electric drill is used for boring holes for the heat flow probe as well as for collecting a core; it has a nominal bit speed of 280 rpm and a percussion

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rate of 2270 bpm. The core is collected in a titanium steel tube that breaks into six sections. The core bit has tungsten-carbide tips. Helical flutes on the outside of the tubes carry excess soil to the surface. Inside diameter of the core sections is 2.04 centimeters. In tests that have been conducted at the NASA Manned Spacecraft Center, only the outermost few millimeters of core have been affected drastically by wall friction.

After drilling is completed, the drill is removed, capped and plugged, then placed in a nylon bag for return. Three of the sections on the Apollo 15 core would not separate; these sections were returned as one section and were broken apart in the laboratory. The uncovered three sections that would not separate (15001 to 15003) were plugged on the lunar surface and taped in the lunar module. The exteriors of the linked core sections were exposed to the atmosphere of the lunar module and command module cabins and had water spots on them (probably caused by sea water splashing into the cabin through an open door after splashdown). The remaining sections were protected by the nylon bag, but the exteriors were exposed to the air in the cabin.

SAMPLING PROCEDURES IN THE LABORATORY

During the preliminary examination of the Apollo 15 samples, the cores were X-rayed and small samples were collected from the bottom of each stem. To obtain a stereo pair, the sections were X-rayed twice by means of a medical X-ray unit in the Lunar Receiving Laboratory. The exposure settings were 93 kilovolts, 100 milliamperes, and 1/5 second, with the X-ray head 0.9 meter above the sample.

Study of the X-radiographs provides a crude indication of grain size and composition of the soil, and the location and orientation of coarser rock fragments (except some breccia fragments, which were transparent at these settings). These images aid in the definition of textural units when the cores are dissected. There is some parallax distortion in the X-radiographs that must be taken into consideration when they are analyzed.

For early allocation to principal investigators and medical staff, 3.75 grams of soil were removed from the bottom of each stem. Included were representative samples collected under red light for thermoluminescence investigations. The entire core cutting and dissection was conducted in a nitrogen atmosphere within stainless steel cabinets that were cleaned to NASA cleaning procedure 2 specifications.

The core was mounted in a cradle, then split in half lengthwise with a milling machine. The split core was removed to another cabinet in a transfer container, where it was mounted on a stage and the upper split tube was removed. The core was then photographed (table I) and described.

The samples then were removed in 0.5-centimeter intervals along the length of the tube, to a depth of approximately two-thirds of the tube diameter. These segments were taken apart with small tweezers, needles, and scoops to get a better overall description of each. If a contact between two stratigraphic units was less than 0.5 centimeter away, the sample was collected up to the contact. Because of slumping, the

location of individual samples was generally within ± 0.5 centimeter. For thermoluminescence studies, representative samples for each stratigraphic unit were collected under a red dark-room light.

The samples were sealed in stainless steel and aluminum or stainless steel and Teflon containers for storage. Allocations to investigators were made from these samples.

After dissection, the remaining one-third of the core was removed to a laminar-flow clean bench where it was impregnated with n-butyl methacrylate to make peels. The peels were kept as a permanent stratigraphic record and as a source of oriented grains from the cores.

DEPTH RELATIONSHIPS AND DRILLING SPEED

According to reference 3, the rotary percussion action of the drill did not significantly disturb the core stratigraphy in premission simulations or on the moon. It is thought that the depth relationships were very close to one-to-one. The core is 242 centimeters long and has a mass of 1333.2 grams. Bulk densities range from 1.62 to 1.93 gm/cm³ (ref. 3).

Variable drilling rates during the boring of the hole for the heat flow experiment and during the collection of the core implied that the regolith at this location consists of a layered sequence. The dissection of the core verified that this was the case.

GENERAL DESCRIPTION OF THE CORE

The identification of stratigraphic units was based on changes in combinations of color, texture, structure, and estimated composition of the coarser rock fragments (table II). Only limited identification of coarser particles with the unaided eye were made during the description and dissection. Petrographic studies were not permitted.

The most common particle type was medium- to dark-grey microbreccia, generally subangular to subrounded, with equant to elongate shapes. Less common were white, feldspar-rich basalt and anorthositic fragments. Near the base of the core, there were medium grey, vesicular and nonvesicular basalt fragments. Details of clast types, and so forth, generally were masked by dust coatings. Black or dark-brown glass droplets and angular glass fragments were present in the soils but were more abundant near the top and bottom of the core.

Several layers at -24.5 centimeters and -83.2 to -94.7 centimeters contained green glass spheres and clastic rocks composed of green glass spheres. This unique glass was present in many of the surface soil samples (ref. 4) and was most abundant at station 7, along the Apennine Front.

Textures ranged from silt-size to pebbly, medium-sand-size soils. The silt-size matrix was ubiquitous and present, to varying degrees, in all layers. All of the

soils were poorly sorted to extremely poorly sorted. The grain size determinations were based on subjective visual and tactile impressions of the soils.

Colors varied from very dark grey (10 YR 3/1) to white (10 YR 8/1). The most common soil color was grey, modified only slightly in value and chroma.

Boundaries between units were generally quite distinct; they could be easily outlined during the dissection. With the exception of less than 1 millimeter of soil along the tube walls, there appeared to be no distortion or smearing of the soil or mixing of layers during the drilling process.

Individual layers ranged from a few millimeters to 13 centimeters thick. A total of 42 major textural units were described within the core. Grading (normal and reverse) of several beds implied that they might have been deposited by turbulent flows, possibly a base-surge type of ejecta cloud. It is also possible that the sorting in an individual layer might have been caused by the pelting of the developing soil surface by micrometeoroids.

CONCLUDING REMARKS

The regolith at the location explored by the crew of Apollo 15 is composed of many thin layers, most of which are probably ejecta from near and far impact events exhibiting a complete spectrum of energies. The regolith at this location is not homogeneous, and no systematic variations are exhibited from bottom to top. This conclusion is tentative, based on only a limited study of the core.

Manned Spacecraft Center
National Aeronautics and Space Administration
Houston, Texas, November 21, 1972
914-40-52-00-72

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3. Carrier, W. David, III; Johnson, Stewart W.; Carrasco, L. H.; and Schmidt, Ralf: Core Sample Depth Relationships: Apollo 14 and 15. Lunar Science III, ed., Carolyn Watkins, 1972, pp. 122-124.
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TABLE I. - NASA PHOTOGRAPHIC COVERAGE OF THE APOLLO 15 DRILL STEM

Section	Predissection	During dissection	Postdissection	Peels	Peel impregnation, etc.
15001	S-72-34095 - 34099 C ^a		S-72-35088 - 35095 B ^b S-72-35151 - 35158 C	S-72-35728 - 35734 B S-72-35684 - 35688 B; RF's ^c stereo S-72-35735 - 35741 C S-72-35659 - 35663 C; RF's stereo	S-72-35159 B S-72-35162 C S-72-35664 C S-72-35669 C
15002	S-72-30757 - 30762 C	S-72-30805 RF S-72-30806 RF S-72-30807 RF	S-72-35079 B S-72-35172 - 35177 B S-72-35179 C S-72-35165 - 35171 C	S-72-35714 - 35720 B S-72-35748 - 35754 C	S-72-35160 B S-72-35163 C S-72-35666 C S-72-35668 C
15003	S-72-44653 - 44660 B S-72-44661 - 44668 C				
15004	S-72-31814 - 31819 C	S-72-32280 RF B	S-72-35080 - 35087 B S-72-35178 C S-72-35180 - 35186 C	S-72-35721 - 35727 B S-72-35742 - 35747 C	S-72-35161 B S-72-35164 C S-72-35665 C S-72-35667 C
15005	S-72-16220 - 16229 C			S-72-19281 - 19296 C	S-72-17893 C
15006	S-72-35647 - 35652 C			S-72-38100 - 38108 B S-72-37945 - 37955 C	S-72-38099 B S-72-38100 B S-72-38101 B

^aThe letter "C" designates a color photograph.

^bThe letter "B" designates a black and white photograph.

^cThe letters "RF" indicate a special photograph of a distinctive rock fragment.

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TABLE II. - DESCRIPTION OF THE APOLLO 15 DRILL CORE

Scale; apparent distance below lunar surface, cm	Permanent unit designations	Temporary unit designations	*Scale; from the top of each metal stem, cm	Sketch of core	Sketch of x-radiograph	Photograph of core	Photograph of a peel from the core	Lithologic Description				
								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
0			0					10 YR 3/1 (Very dark grey)	Fine sand-bearing silt-size soil. Contains about 1-3% particles greater than 1 mm in diameter. Poorly sorted. Coarser than Unit V.	Subangular to angular blocky structures, 3-4 mm long. Upper 2.5 cm has been somewhat disrupted by the plug. There are some 1-2 mm long cohesive aggregates from 4-5 cm below the top of the core.	1. Black or dark brown glass fragments and spheres, generally less than 1 mm in diameter. 2. Medium to light grey, subrounded to subangular (breccia?) fragments. 3. Subrounded to subangular, white (anorthosite?) fragments.	None
1	42	006-VI	1									
2			2									
3			3									
4			4									
5			5									
6	41	006-V	6									
7			7									
8			8					10 YR 3/1-4/1 (Dark grey to very dark grey)	Sand-bearing, silt-size soil. Only a trace of particles greater than 1 mm in diameter.	Less coherent than Units VI and IV. Weak, subangular blocky structures, 3-5 mm long.	1. Black to dark brown glass spheres and fragments. 2. Light grey to grey (breccia?) fragments. 3. White (anorthosite?) fragments.	None
9			9									
10			10									
11	40	006-IV	11									
12			12									
13			13									
14			14					10 YR 4/1 (Dark grey)	Sand-bearing silt-size to silt-size soil; poorly sorted. Trace to 3% particles greater than 1 mm in diameter.	Weakly coherent with irregular weak prismatic (3 x 10 mm) structures. Some light grey cohesive aggregate (clods).	1. Dark brown to black glass agglutinates and fragments (some are vesicular). 2. Light to dark grey (breccia?) fragments. 3. White (anorthositic?) rock fragments.	None
15			15									
16	39	006-III	16									
17			17									
18			18									
19			19					10 YR 5/1 (Grey)	Silt-size soil; poorly sorted. Trace to 5% particles greater than 1 mm in diameter.	Weak, subangular blocky to prismatic structures (2-7 mm long) strong tendency to form slabs along wall of open core. 3 to 10% of the soil consists of lighter grey, irregular, blocky patches of soil. These "patches" have nearly the same coherence as the surrounding darker grey soil.	1. Light grey to grey breccia fragments (subrounded). 2. Black to dark brown glass fragments and spheres. 3. Fragile, powdery white fragments; possibly anorthosite. 4. Green glass spheres at 24.5 cm below the top of the core.	None
20			20									
21	38	006-II	21									
22			22									
23			23									
24			24									
25			25									
26	37	006-I	26					10 YR 5/1 (Grey)	Sand-bearing silt-size soil; very poorly sorted. 0 to 5% (near base) particles greater than 1 mm in diameter; may be graded.	Subangular, blocky structures, 1-5 mm long. Weakly coherent.	1. Fine sand size to 2 mm white (anorthositic?) fragments. 2. Light grey (breccia?) fragments. 3. Trace of black glass fragment.	None
27			27									
28			28									
29			29									
30			30									
31	37	006-I	31									
32			32									
33			33					10 YR 4/1-5/1 (Grey to dark grey) for Subunit B 10 YR 5/1 (Grey) for Subunit A	Sand-bearing silt-size soil; 0 to 2% particles greater than 1 mm long. (Very fine grained).	Weakly coherent, subangular, blocky structures (less than 2 mm in diameter) Subunit B is less coherent than Subunit A. Some 1-2 mm diameter cohesive aggregates (clods).	1. Dark brown to black glass fragments and droplets. 2. Light grey, subangular (breccia?) fragments. 3. White (anorthositic?) fragments.	B. Less coherent, darker grey. A. More coherent, lighter grey.
34			34									
35			35									
36			36									
37			37									
38			38									
39			39									
40			40									

*The individual sample locations are based on this scale.
**Munsell Color Co., Inc., Munsell Soil Color Charts, 1954 ed., Baltimore, Md.

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TABLE II. - DESCRIPTION OF THE APOLLO 15 DRILL CORE - Continued

Scale; apparent distance below lunar surface, cm	Permanent unit designations	Temporary unit designations	*Scale; from the top of each metal stem, cm	Sketch of core	Sketch of x-radiograph	Photograph of core	Photograph of a peel from the core	Lithologic Description				
								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
40	37	006-I-A	40	Cont				See previous page for description of unit.				
41			41									
42			42									
43	36	005-VIII	3									
44			4					10 YR 5/1-6/1 grey	Sandy silt-size soil	Forms coarse friable aggregates. Slumped along edges.	Subunit 4 has a lense of small white and grey soil aggregates. Some grey microbreccia.	None
45			5									
46			6									
47	35	005-VII	7									
48			8					10 YR 4/1-4/2 dark grey to dark greyish brown	Silt-size; very uniform. Some sand-size particles at base; may be graded.	Less cohesive than the units above; few aggregates.		None
49	34	005-VI	9									
50			10									
51			C 11					10 YR 5/1 grey	Fine sand size to silt size; containing 5-10% coarse sand to granule-size particles.	Slumps along edge. Soil breaks into loosely coherent aggregates.	Coarser fragments appear to be mostly grey to white microbreccias.	C. Contains white aggregates. (See textural description). B. Lens of granule-size rock fragments and abundant white, friable aggregates. A. Similar to Subunit C.
52			B 12									
53	33	005-V	A 13									
54			G 14									
55			F 15									
56			E 16									
57			D 17					10 YR 4/1 (dark grey) near the top; grades down to 10 YR 5/1-5/2 (grey to greyish brown)	Silt-size soil with 2-5% coarse sand or granule size particles.	Consist of coherent aggregates; most are the same color as the matrix, but some are lighter. Probably a coherent layer broken by the drill.	Light grey microbreccias and crystal fragments.	G. Slightly finer grained than the over and under- lying units. F. Slightly lighter color than the over and under- lying units. E. Abundant white or light grey coherent aggregates 0.5-1.0 mm long. D. Coarser than B. Contains abundant coherent aggre- gates which are the same color and texture at the matrix. C. Lens containing about 10% particles > 1 mm. B. Finer grained than C. A. Slightly darker than Subunit B.
58			C 18									
59			B 19									
60			A 20									
61			D 21									
62			C 22									
63	32	005-IV	B 23									
64			A 24									
65	31	005-III	25									
66			26									
67			27					10 YR 4/1 (dark grey)	15% coarse sand to granule-size particles in silty matrix.			None
68			28									
69			29									
70			C 30									
71			B 31									
72			A 32									
73			D 33					10 YR 5/1 (grey) near the top, grading down to 10 YR 6/1 (grey) near the base.	10-20% particles > 2 mm in a silt-size matrix.	Forms cohesive aggregates (clods). This was probably a very cohesive layer; broken up by the drilling.	Microbreccia fragments in coarse fraction.	C. Contains a lens of coarse- sand and granule-size particles. B. Large microbreccia fragments. A. Fractured unit; lighter than Subunit B.
74			C 34									
75			B 35									
76			A 36									
77			D 37									
78			C 38									
79			B 39									
80			A 40									

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								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
80	31	005-II	40					10 YR 5/2 greyish brown	Coarse sandy soil.	Uniform, unbroken layer.	Abundant brown glass droplets and agglutinates.	None
81			41					10 YR 6/1 grey	Granule and coarse sand bearing silt size soil.	Cohesive; forming small clumps up to 5 mm long.	(1) abundant microbreccia fragments	None
82	30	005-I	42									
83			43									
84	29	004-IV	2					10 YR 5/1 to 10 YR 5/1 grey	Poorly sorted granule-bearing silt-size soil. Coarser fragments vary from 10% near the top to 0 near the base of the bed; it is reversely graded.	Moderate slumping; forming 4-5-mm long clumps.	1. Light grey to dark grey breccias; some with dark brown glass coatings. 2. Clastic rocks consisting of green glass spheres and isolated spheres. 3. Trace of small, powdery white "anorthositic" fragments.	None
85			3									
86			4									
87			5									
88			6									
89			7									
90			8									
91			9									
92			10									
93			11									
94			12									
95	28	004-III	13					10 YR 5/2 greyish brown (patches of lighter grey soil)	0-5% granule to coarse sand-size particles in silt-size soil.	Moderate slumping; forming 4-5-mm long clumps.	1. White, powdery "anorthosite" fragments. 2. Light grey breccia fragments; a few have dark grey-brown glass coatings. 3. White aggregates of "clods".	None
96			14									
97			15									
98			16									
99			17									
100			18									
101	27	004-II	19					10 YR 5/1 grey	Coarse sand to granule-bearing fine sand-size soil. Possible reversely graded.	Forms very fragile blocks, 2 x 4 to 10 mm long; slumps easily.	1. Light to dark grey breccia fragments. 2. Grey and black angular glass fragments. 3. Trace of green glass.	None
102			20									
103			21									
104			22									
105			23									
106			24									
107	26	004-I	25					10 YR 5/1 grey	Moderately well-sorted silt-size material. 1-3% of the volume is composed of fragments greater than 1 mm long.	Forms coherent blocks 2-5 mm long. There was little collapse when the core was opened.	1. Light grey breccia fragments. 2. Trace of dark brown and grey glass fragments.	None
108			26									
109			27									
110			28									
111			29									
112			30									
113			31									
114			32									
115			33									
116			34									
117			35									
118			36									
119			37									
120			38									

*The individual sample locations are based on this scale.
**Munsell Color Co., Inc., Munsell Soil Color Charts, 1954 ed., Baltimore, Md.

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No peel was made of Core 15003.

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								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
160	19	003-I	38									
161			39									
162			40									
163			41									
164			42									
165			43	Void								
166		002-XI	3									
167			4									
168	18	002-X	5									
169			6									
170			7									
171			8									
172			9									
173	17	002-IX	10									
174			11									
175			12									
176			13									
177	16	002-VIII	B 14									
178			15									
179			16									
180			A 17									
181			18									
182			19									
183	15	002-VII	20									
184			21									
185			22									
186	14	002-VI	23									
187			24									
188			25									
189			26									
190	13	002-V	C 27									
191			28									
192			F 29									
193			30									
194	12	002-IV	31									
195			E 32									
196			33									
197			D 34									
198			C 35									
199			A 36									
200	11	002-III	37									

*The individual sample locations are based on this scale.

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See next page for a description of this unit.





















TABLE II. - DESCRIPTION OF THE APOLLO 15 DRILL CORE - Continued

Scale; apparent distance below lunar surface, cm	Permanent unit designations	Temporary unit designations	*Scale; from the top of each metal stem, cm	Sketch of core	Sketch of x-radiograph	Photograph of core	Photograph of a peel from the core	Lithologic Description				
								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
200	11	002-III	37					Grey (5 Y 5/1)	Silt size; moderately well sorted.	No cohesive aggregates.	None	None
201	10	002-II	38					Grey (5 Y 5/1)	Coarse sand and granule-bearing silt-size soil.	No cohesive aggregates.	Too dusty for identification.	None
202			39									
203			40									
204			41									
205	9	002-I	42					Grey (5 Y 5/1)	Granule-bearing fine silt-size soil.	Some white cohesive aggregates. Prismatic fracture pattern.	1. White "anorthosite" 2. Feldspar crystals.	B. Abundant black glass droplets. A. Some anorthositic fragments; about 20% fragments > 1 mm.
206			1									
207			2									
208			3									
209	8	001-VIII	4					10 YR 7/1 light grey	Silt-size soil; moderately well sorted.	No cohesive aggregates.	1. Light grey to white breccia fragments. 2. Some black glass fragments.	None
210			5					10 YR 7/1 light grey	Granule-bearing silt-size soil.	≈ 2% cohesive, grey aggregates.	1. Vesicular black glass fragments. 2. Light grey breccias.	None
211	7	001-VII	6									
212			7									
213	6	001-VI	8					10 YR 7/1 light grey	Silt-size soil.	Some brown and white cohesive aggregates.	1. Light grey breccia fragments. 2. Black glass fragments and droplets.	None
214			9									
215	5	001-V	10					10 YR 7/1 light grey	Pebbly silt-size soil.	Abundant white and grey cohesive aggregates.	1. Grey breccia fragments. 2. Black glass fragments. 3. Basalt(?) fragments.	C. 20-30% granule-size fragments; abundant black glass. B. > 20% granule-size fragments; abundant clods. A. ≈ 5% granule-size fragments; abundant clods.
216			11									
217	4	001-IV	12					10 YR 8/1 white	Silty fine sand-size soil.	Distinct fine (0.5 mm) laminae. Less coherent than Unit V.	1. Abundant black glass droplets and fragments. 2. Grey breccia fragments.	B. Slightly coarser and more black glass than Unit A. A. Abundant coherent aggregates (clods).
218			13									
219			14									
220			15									
221			16									
222			17									
223	3	001-III	18									
224			19									
225			20									
226			21									
227			22									
228			23									
229			24									
230			25									
231			26									
232	2	001-II	27					10 YR 7/1 light grey	Granule-bearing sandy silt-size soil.	Angular, 1-2 mm blocky structures, with thin, long prisms along the interior of the drill stem.	1. Light grey breccia fragments. 2. Medium grey, aphanitic basalt fragments. 3. Trace of black glass.	F. Possible graded bed, with 10-40% fragments > 1 mm. E. Less basalt(?) fragments than Unit E. D. Silt-size sand. C. 25% fragments > 1 mm. B. Silt-size soil. A. 12-50% fragments (light-grey breccias) > 1 mm.
233			28									
234			29									
235			30									
236			31									
237			32									
238	1	001-I	33					10 YR 7/1 light grey	Silt-size soil.	A few clods present.	1. Light grey breccia fragments. 2. Vesicular grey basalt. 3. Dark brown glass.	F. Silt-size soil. E. 10-20% particles > 1 mm, in a silty matrix. D. Silt-size soil. C. Some aggregates (clods) present. B. Silt-size soil. A. Silt size, with some > 1 mm diameter clods.
239			34									
240			35									

*The individual sample locations are based on this scale.
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TABLE II. - DESCRIPTION OF THE APOLLO 15 DRILL CORE - Concluded

Scale; apparent distance below lunar surface, cm	Permanent unit designations	Temporary unit designations	*Scale; from the top of each metal stem, cm	Sketch of core	Sketch of x-radiograph	Photograph of core	Photograph of a peel from the core	Lithologic Description				
								**Color	Texture	Structure	Composition of larger rock fragments	Subunits
240	1	001- I	B 35					See previous page for description.				
241			A 36									
242												
243												
244												
245												
246												
247												
248												

*The individual sample locations are based on this scale.
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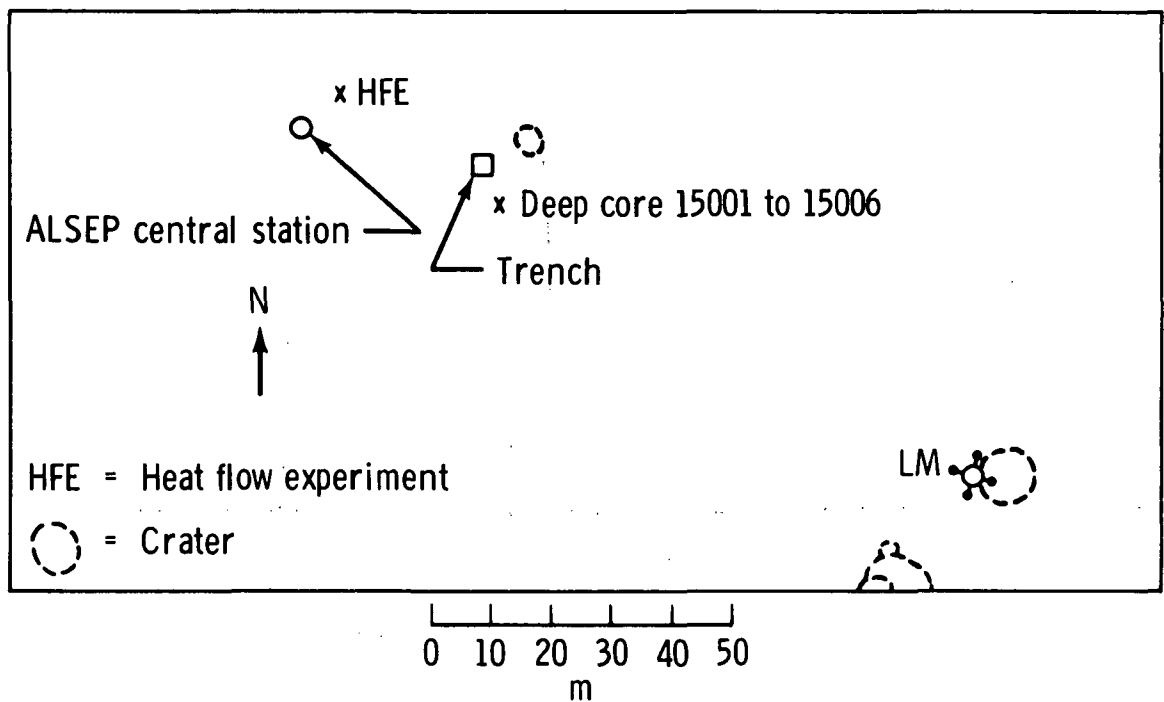


Figure 1.- Sketch map of the ALSEP-lunar module site area.